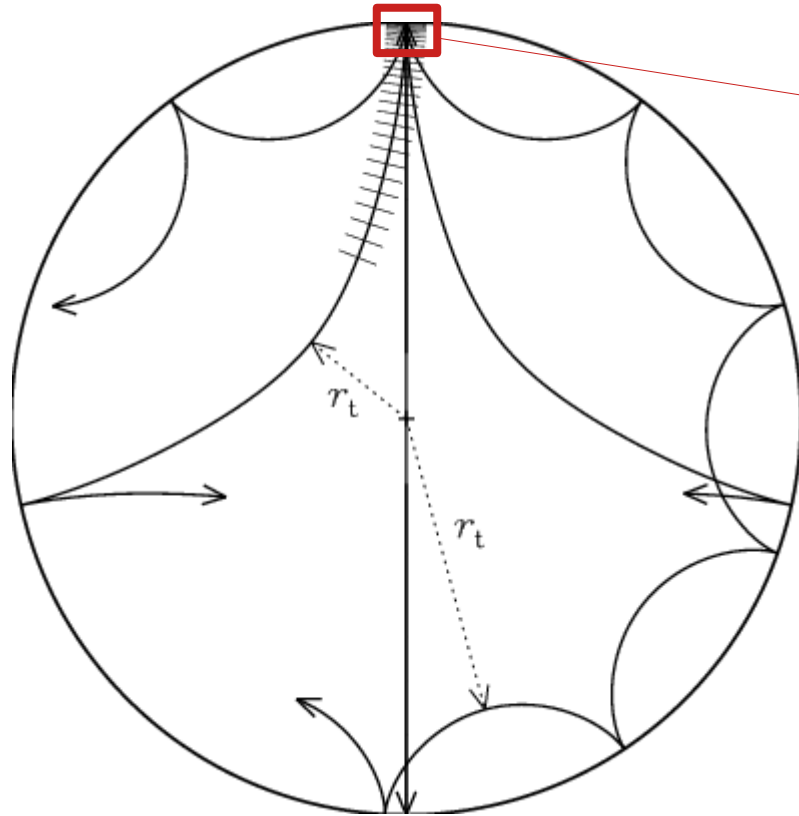


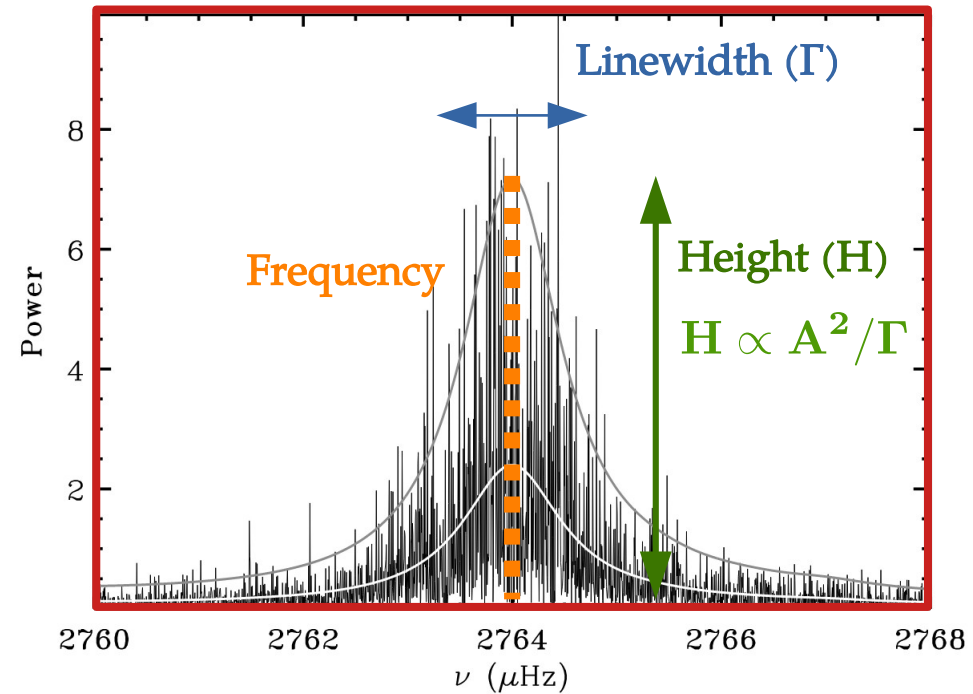
PDF Lagrangian models of turbulence for the modelling of convection/oscillation coupling

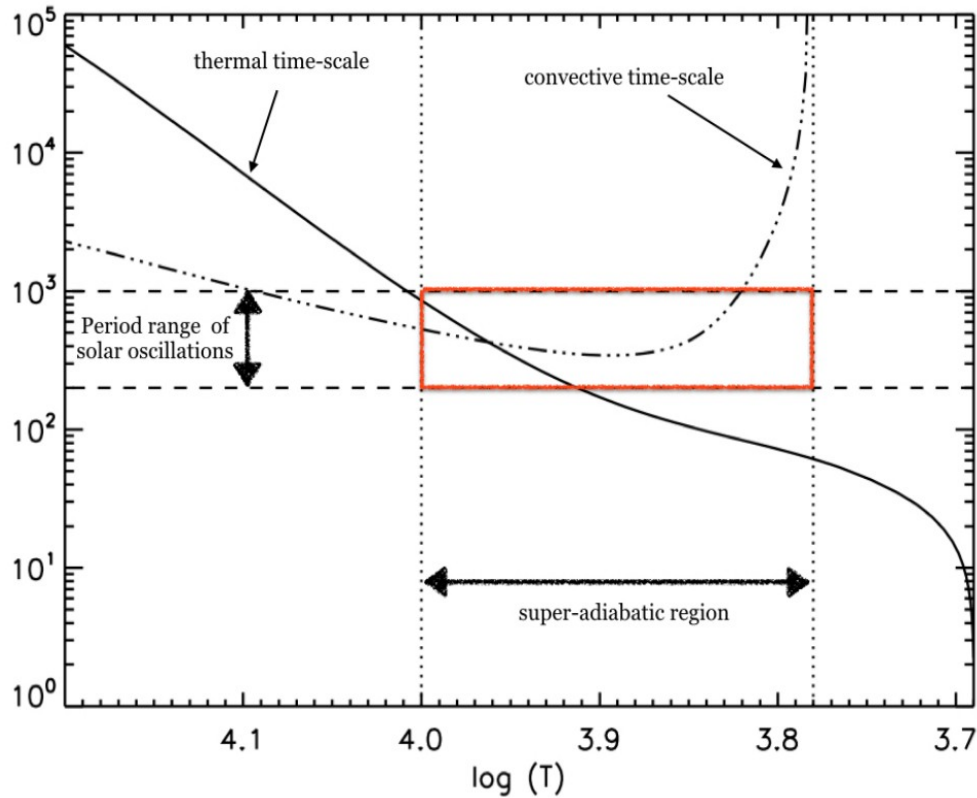
Jordan Philidet
(sur l'aimable invitation de Kevin Belkacem)

P-modes: propagation in turbulent medium
→ effect on the frequency
→ driving
→ sink of energy = damping



Localised near the surface





- Coupling near surface because

- coincidence time scales

$$\tau_{\text{turb}} \sim \tau_{\text{th}} \sim \Pi \sim 5 \text{ minutes}$$

- coincidence spatial scales

$$\lambda \sim l_{\text{conv}} \sim 1 \text{ Mm}$$

- Good news: seismic observables can be used to constrain turbulent convection

- Bad news: theoretical modelling is very complicated



- Today: done with MLT, Reynolds-stress models. Also valuable insight from simulations
- But not ideal: full effect of turbulent cascade can't be included → turbulent dissipation?

- Core idea → represent turbulent flow with set of individual fluid parcels
- Turbulent flow → particle evolution follows stochastic differential equations

$$df^* = \underbrace{a(f^*, t) dt}_{\text{Deterministic}} + \underbrace{b(f^*, t) dW}_{\text{Random}}$$

- ω_t = turbulent frequency → inverse of turbulent energy decay rate
- C_0 = Kolmogorov constant → universal value
- G_{ij} = drift tensor → given by equivalent Reynolds-stress model

- Simple model

$$dX_i^* = U_i^* dt$$

$$dU_i^* = \underbrace{-\frac{1}{\bar{\rho}} \frac{\partial \bar{p}}{\partial x_i} dt + g_i}_{\text{Mean pressure force + gravity}} + \underbrace{G_{ij}(U_j^* - \bar{U}_j) dt + \sqrt{C_0 \widetilde{u_i'' u_i''}} \omega_t dW_i}_{\text{Fluctuating pressure force + buoyancy + turbulent dissipation}}$$

Mean pressure force + gravity

Fluctuating pressure force + buoyancy + turbulent dissipation

But: necessary to evaluate the mean fields at the particle position

How to estimate the mean fields ?

In general, extraction from DNS/LES

Here, impossible because

- we need
- that would decouple oscillations and turbulence

Kernel $W(r,h)$

Particle of interest

Neighbor particles

h

Solution: estimation directly from fluid parcels

→

$$\langle Q(\mathbf{x}, t) \rangle_{N,h} \equiv \frac{V}{N} \sum_{i=1}^N W \left[\mathbf{x}^{(i)}(t) - \mathbf{x}, h \right] Q \left(U^{(i)} \right)$$

Example for density

$$\bar{\rho}(\mathbf{x}, t) = \frac{V}{N} \sum_{i=1}^N W \left[\mathbf{x}^{(i)}(t) - \mathbf{x}, h \right]$$

Credits: Yi et al. (2017)

Linearisation:

$$\begin{aligned} \frac{\partial \boldsymbol{\xi}_{\text{osc}}}{\partial t} - \mathbf{u}_{\text{osc}} & - (\boldsymbol{\xi}_{\text{osc}} \cdot \nabla) \mathbf{u}_t = \mathbf{0} \\ \frac{\partial \mathbf{u}_{\text{osc}}}{\partial t} - \mathbf{L}_1^d & - \mathbf{L}_1^s = \mathbf{L}_0 \end{aligned}$$

Deterministic propagation
→ p-modes without turbulence

Stochastic perturbation of wave propagation
→ damping + surface effects

Stochastic inhomogeneous term
→ driving

Philidet et al. (2021)

Expansion over normal modes:

$$|z(t)\rangle = \sum_i A_i(t) \exp^{j(\omega_i t + \Phi_i(t))} |z_i\rangle$$

Unperturbed eigenfunction
Perturbed mode amplitude
Perturbed mode phase

Time-dependent perturbation → linear coupling between modes
But for now, let us focus on single mode case

Stochastic perturbation has memory

→ effect on mean amplitude and mean phase must be accounted for

→ simplified amplitude equation formalism (Stratonovitch 1965, Buchler et al. 1993)

$$dA = \left(A(\kappa + \text{Re}(\alpha_3)) + \frac{\text{Re}(\alpha_1)}{2A} \right) dt + \left(A^2 (2\alpha_2^R + \text{Re}(\alpha_3)) + \text{Re}(\alpha_1) \right)^{1/2} dW_A$$
$$d\Phi = \text{Im}(\alpha_3) dt + \left(\frac{1}{A^2} \text{Re}(\alpha_1) + 2\alpha_2^I + \text{Re}(\alpha_3) \right)^{1/2} dW_\Phi$$

Philidet et al. (2022)

Physics contained in complex $\alpha_i \rightarrow$ autocorrelation spectrum of stochastic perturbation

Mean energy $E_m = \langle A^2 \rangle$ and
mean phase $\Phi_m = \langle \Phi \rangle$

$$\frac{dE_m}{dt} = -2\eta E_m + \mathcal{P}$$

Damping rate

Excitation rate

$$\frac{d\Phi_m}{dt} = \delta\omega$$

Modal surface effects

$$\mathcal{P} = \frac{1}{2\mathcal{I}} \int d^3\mathbf{X} \rho_0^2 k_j k_l \operatorname{Re} \left(\xi_{\text{osc},i} \xi_{\text{osc},k}^* \phi_{ijkl}^{(4b)}(\mathbf{k}, \omega) \right)$$

Depends on equilibrium structure + mode structure
Depends on spectrum of turbulent fields

$$\eta = -\frac{1}{4\mathcal{I}^2} \int d^3\mathbf{X} \rho_0^2 \operatorname{Re} \left(F_i^{(1)} F_j^{(1)*} \phi_{ij}^{(2)}(2\mathbf{k}, 2\omega) + F_{ij}^{(2)} F_{kl}^{(2)*} \phi_{ijkl}^{(4b)}(2\mathbf{k}, 2\omega) + F_{ijm}^{(3a)} F_{kln}^{(3a)*} \phi_{ijklmn}^{(4c)}(2\mathbf{k}, 2\omega) \right. \\ \left. + 2F_i^{(1)} F_{jkl}^{(3b)*} \phi_{ijkl}^{(4a)}(2\mathbf{k}, 2\omega) + 2F_{ijm}^{(3a)} F_{kl}^{(2)*} \phi_{ijklm}^{(4d)}(2\mathbf{k}, 2\omega) \right)$$

$$\delta\omega = \frac{1}{4\mathcal{I}^2} \int d^3\mathbf{X} \rho_0^2 \operatorname{Im} \left(F_i^{(1)} F_j^{(1)*} \phi_{ij}^{(2)}(2\mathbf{k}, 2\omega) + F_{ij}^{(2)} F_{kl}^{(2)*} \phi_{ijkl}^{(4b)}(2\mathbf{k}, 2\omega) + F_{ijm}^{(3a)} F_{kln}^{(3a)*} \phi_{ijklmn}^{(4c)}(2\mathbf{k}, 2\omega) \right)$$

Example
turbulent spectra

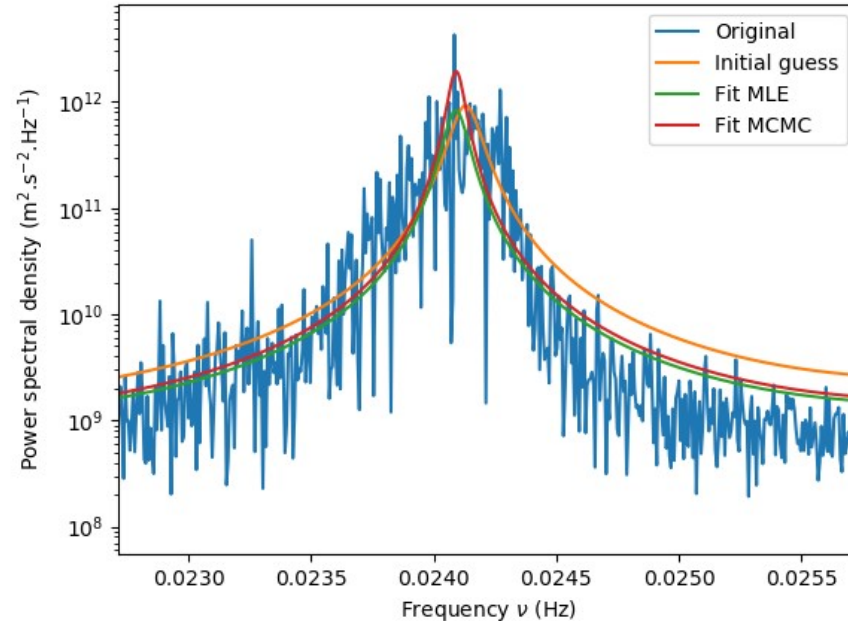
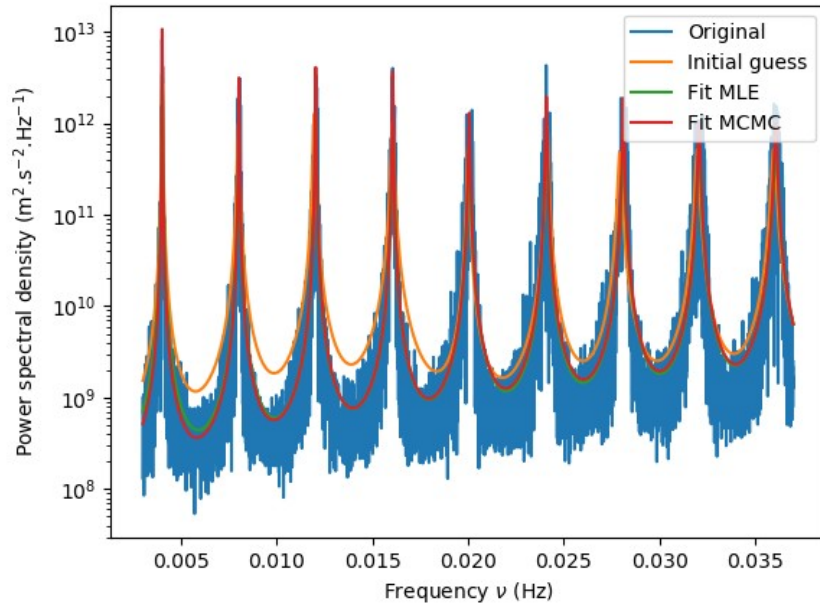
$$\phi_{ij}^2(\mathbf{k}, \omega) \equiv \int_{-\infty}^0 d\tau \int d^3\delta\mathbf{x} \langle u_{t,i}(\mathbf{X}, t) u_{t,j}(\mathbf{X} + \delta\mathbf{x}, t + \tau) \rangle \exp^{j(\mathbf{k} \cdot \delta\mathbf{x} + \omega\tau)}$$

- Consistent with known expression of excitation rate (e.g Samadi & Goupil 2001)
- Damping rate and modal surface effects are real and imaginary part of same complex quantity
- Allows for the injection of any prescription for turbulent velocity spectrum

That was for very simple Lagrangian model: for more complete model, need to go numerical

- Very simple to implement: no spatial grid or scheme needed
- Time scheme adapted to stochastic nature of equations
- Output of simulation = kernel averages computed from set of particles

For the moment, 1D implementation: test theoretical formalism



2D and 3D implementations are under way. Investigation of:

- Coupling with gravity waves
- Lagrangian formalism → perfectly adapted to Lagrangian tracers → transport (chemical, angular momentum, ...)

Still very much pending though

Danke sehr!!